

An Introduction to the HDD, modelling, detection and decoding for magnetic recording channels

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Outline

 Overview of the Hard Disk Drive Fundamental problem for magnetic recording HDD current and future technologies

- 2 Magnetic Recording Media
- 3 Magnetic Recording Head Writer Reader
- 4 Servo
- 5 Magnetic Recording Channel (MRC)
 - The Encoder Channel Model The Equalizer The iterative detector The joint Viterbi detector/decoder
- 6 Future HDD Technologies



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Overview of the Hard Disk Drive



VCM (Voice Coil Motor): Actuator for positioning the head

Arm: Rigid mounting for the head

HDI: Head disk interface

Servo: Control scheme for positioning the head over a $\sim 40-50 nm$ track

Mechanics: Managing vibration, resonant freq, airflow, etc... for the servo

Coding and Signal Processing: Encoding the user data so that it can survive the passage through the channel and detecting/decoding the data on readback. Our work also involves creating good models for the recording channel.



Fundamental problem for magnetic recording

Defining terms

Magnetic Anisotropy:

- Isotropic: The property of being independent of direction.
- · Anisotropic: Not isotropic, dependent on the direction
- Magnetic anisotropy: A property of a magnetic moment or material that prefers to orient in a certain direction (the easy axis direction)
- Magnetic anisotropy field (symbol: *H_k*, units:kOe): A non-real magnetic field that can be imagined to be holding the moment in the easy-axis direction.
 The moment behaves as if such a field exists, though it does not.
- Magnetic crystalline anisotropy constant (symbol: K_u, units: J/m³): The amount of energy needed to cause a volume of the magnetic material to flip

Magnetic anisotropy is the property that enables recording.



Fundamental problem for magnetic recording

The superparamagnetic limit and the media trilemma

3 different ways different ways (trade off) in which AD growth is halted:

- 1 Grains kept large→run out of SNR
- 2 Grains shrunk:
 - Big K_u → media unwriteable
 - Small K_u → media thermally unstable

Media SNR ~ log(N)

where N=number of grains/bit

Higher Density \rightarrow smaller bits smaller bits \rightarrow smaller grains

Thermal Stability: $\frac{K_u V}{k_B T} \ge 60$

Small grains become thermally unstable: Random variations due to temperature have nonzero probability to flip grains.

Writeability:

Write head field is limited. Media must have low enough K_u so that it can be magnetized by the writer.



HDD Areal density trend



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HDD current and future technologies





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Magnetic Recording Media: Hysteresis loops

- Hysteresis loops plot
 - · Applied field on the horizontal axis vs
 - Media magnetization on the vertical axis
- Field is applied perpendicular to the medium.
- Interesting hysteresis properties of the medium H_n Nucleation field The field required to nucleate the switching H_c Coercive field The field required to switch 50% of the media H_s Saturation field field required to saturate the media M_s Saturation magnetization All moments aligned maximally M_r Remnant magnetization The remaining moment when the field is removed S Squareness $S = H_n/H_s$



Magnetic Recording Media

Defining terms

Media Noise:

- Media noise is noise generated by the granular nature of the medium.
- Transitions written on the medium *must* follow the grain boundaries
 - Larger grains \rightarrow more deviation in transition \rightarrow more noise
 - Smaller grains \rightarrow less deviation in transition \rightarrow less noise
- Media noise manifests as transition jitter in the readback signal.
- · Errors tend to be located around the transitions.

Media noise is the dominant noise source in magnetic recording channels.

Exchange coupling:

- Exchange coupling is a short-distance (nm-range) effect that exists between adjacent grains.
- Exchange coupled grains exert a magnetic force (field) on each other that tends to align the grains in the same direction.
- The stronger the exchange coupling, the more the grains will align.
- Exchange coupling causes grains to clump together into larger clusters.

We would like to have well isolated (exchange decoupled) grains.



Magnetic Recording Media: Issues

Desirable properties of magnetic recording media:

- Smaller grains \rightarrow less media noise \rightarrow higher densities
- Smaller grain-size variation
 - Variation in grain-size \rightarrow more media noise
- Tighter magnetic property distributions
 - Tighter distributions in the magnetic properties → sharper transitions on the medium.
- Engineering multiple layered media
 - Exchange-coupled layers → more thermal stability while maintaining writeability.

 μ -mag simulation hysteresis loops varying each of the parameters in this table are shown

	nominal value		deviated value	
	μ	σ	μ	σ
<i>K_u</i> (J/m ³)	3.6e5	3%	3.6e5	13%
<i>M</i> _s (A/m)	4.8e5	3%	4.8e5	13%
A_x (J/m)	3e-12	3%	3e-11	3%
ez	0°	1.7°	5°	1.7°



Varying σ_{Ku} Deviated $\sigma_{Ku} = 13\%$

Nominal $\sigma_{Ku} = 3\%$



Varying σ_{Ms} Deviated $\sigma_{Ms} = 13\%$

Nominal σ_{Ms} = 3%



Varying A_x

Deviated $A_x = 3e-11$

Nominal $A_x = 3e-12$



Varying *e*_z

Deviated $e_z = 5^{\circ}$

Nominal $e_z = 0^\circ$



Multi-layer media



- Multi-layer media (aka exchange coupled composite media, spring coupled media) provides more thermal stability for the same writeability.
- Each of the multiple layers have different *K*_u values and switch at different fields.
- The ECL layer controls the amount of coupling between layers:
- Soft layers help the harder layers to switch.
- Hard layers provide thermal stability to the soft layers.
- Commercial media today is multi-layered.



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Magnetic Writer



- · Magnetic write head driven by the write current
 - · Write current modulated by the channel bits
- · Longitudinal recording (left)
 - Previous generation technology
 - · Used the fringe field from the head gap to write.
- Perpendicular recording (right)
 - · Perpendicular recording includes a soft underlayer (SUL)
 - Provides return path for the magnetic flux
 - Effectively putting the media in the write-gap
 - Doubles the amount of field that can be generated



Magnetic Writer

- Magnetic writer when viewed from the ABS is trapezoidal in shape.
- · Writer conducts the magnetic flux to the pole tip
- Trailing shield increases the trailing write-field gradient
 - · Prevents overwriting of previously written bit
- · Side shields increase the side write-field gradient
 - · Prevents overwriting of adjacent tracks









Magnetic Write fields

- Simulated magnetic fields generated by the writer are shown.
- 3 Components (Hx, Hy, Hz) for 1D slice of field in each direction
 - Cross-track Profile

 \Rightarrow

1

ll

- 2D Profile
- Down-track Profile
- In-plane components help in the switching.



Down-track profile

Cross-track profile





2D profile



Magnetic writing simulation

Parame	Value	
Grain size	(nm)	9
σ_{gs}		16%
$\tilde{K_u}$	(J/m ³)	240e3
Х $\sigma_{\it Ku}$		3%
Ms	(A/m)	480e3
σ_{Ms}		3%
H_k	(kOe)	10
σ_{Hk}		4.26%
ez		0 °
$\sigma_{\it ez}$		1.7°
velocity	(ms ⁻¹)	20
BL	(nm)	14



Magnetic Reader



SEM x-section image

- Reader technology evolution:
 - 1 AMR Anisotropic magnetoresistive
 - 2 GMR Giant magnetoresistive
- Magneto-resistive: The property of materials to change resistance in the presence of a magnetic field.



Magnetic Reader



- Ambient magnetic field ⇒ free-layer to rotate
- \rightarrow resistance change of the device.
- Pass current through to measure R
- · Reader is sensitive to a region of the magnetic medium
- Characterized by read head sensitivity (RHS) function



Model for the Read head sensitivity (RHS) function





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Servo

- The head needs to be positioned over and follow a ~ 40nm wide track
- Disturbance sources include:
 - Air-flow
 - The Motor
 - The VCM
 - External shock and vibe.
- Servo wedges run radially from ID to OD
- 200-300 servo wedges per revolution
- Head can deduce its position when reading a servo wedge
- Servo control algorithm puts the head "on track"
- Between wedges, flying "blind"



Reverse Engineering the Servo Wedge



Preamble Preamble is used to synchronize/initialize loops

- SAM Sector address mark is an index identifying the current sector
- Burst This gives the information of the cross-track position
- Gray A gray-code that is probably used as a track-index
- Gap A gap between the servo sector and data sector
- Pre Some known pre-data
- Data The encoded data written onto the medium



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Magnetic Recording Channel (MRC)



- The magnetic recording channel (MRC) has many similarities to the communications channel:
 - · The comms channel transports data from here to there
 - · The MRC transports data from now to later
 - Both distort the signal and introduce noise
- Main distortions on the MRC
 - Intersymbol interference (ISI)
 - Media noise
 - Non-linear transition shifts
- Current MRC use LDPC codes and pattern dependent noise predictive detection



The Encoder



- LDPC Block Codes are used in the MRC.
 - · Block codes have a generator matrix G.
 - And parity check matrix H.
- **G** is $K \times N$. Generates the codeword during encoding
 - **c**= $\mathbf{u} \times \mathbf{G}$. **c** is $1 \times N$ codeword, **u** is $1 \times K$ userword.
 - Systematic codes effectively M = N K parity bits to the user word
- **H** is $M \times N$. Checks whether a bit sequence is a legal codeword.
 - · LDPC codes have H matrices with a low proportion of 1's.
 - MRC's use quasi-cyclic codes ⇒ reduce memory requirement



The Encoder



• The code breaks up GF(2^N) into 2 spaces:

The space of legal codewords

• $\mathbf{c} = \mathbf{u}\mathbf{G}$ and $\mathbf{c}\mathbf{H}^T = \mathbf{0}$.

2 The space of non-codewords

- cH^{*T*} ≠ 0.
- *d_{min}*: minimum distance btw codewords
- Desirable qualities for the code:
 - Large d_{min}
 - No short cycles (large girth)
 - Longer CWL $\Rightarrow \uparrow$ performance
 - Low/no error floor





The Encoder



- Error floors are the bugbear of the coding designers
- Error floors can manifest from different sources
- · Error floor of interest here: from the code
 - Subject of significant research effort
 - Caused by trapping sets in the code FG
- MRC need to deliver error rates of 10⁻¹⁴
 - · Difficult to check for error-floors
- HDD industry switching from 512byte→4096byte.
 - Longer CWL \Rightarrow better performance.





The Channel Model



- To realistically reproduce waveforms for channel simulations
- Simple model:
 - · Additive white Gaussian noise (AWGN) Model
 - Everybody's first time model
- Intermediate model
 - Jitter channel model
 - · Captures more realistically the media noise characteristic
 - · Divorced from the physics of recording
- Comprehensive MRC Model
 - µ-mag simulation and GFP model.



Micromagnetic (μ -mag) simulations

- μ -mag simulations involve solving the Landau-Liftshitz Gilbert (LLG) equations numerically

• LLG:
$$\frac{d\mathbf{M}}{dt} = \mathbf{M} \times \mathbf{H}_{tot} - \alpha \mathbf{M} \times (\mathbf{M} \times \mathbf{H}_{tot})$$

- M is the magnetization vector
- **H**_{tot} is the total field experienced by **M**.
- α is a phenomenological damping parameter that helps to match the model to experimental results
- The LLG describes the behavior of a magnetic particle ${\bf M}$ in the presence of a magnetic field ${\bf H}_{tot}.$
- In μ -mag simulations **H**_{tot} is the sum of:
 - **H**_k: The anisotropy field
 - **H**_m: The magnetization from the surrounding grains
 - H_e: The external field from the head
 - **H**_x: The exchange field from exchange coupled grains (nearest neighbors)
- Solving (1) numerically predicts the magnetization behavior



(1)

The Grain-flipping probability (GFP) Model

• μ-mag simulations produce grain magnetizations that we can sample at the end of each bit:



- · The GFP model keeps count of:
 - The grains in the region that could flip (denominator array)
 - · Grains magnetized opposite to the applied field can flip
 - The grains in the region that do flip (numerator array)




The Grain-flipping probability (GFP) Model

- GFP model ties channel simulations to physics of recording $_{H^{k}g^{SMs}e^{Z}}$



- GFP channel model flips grain using a RNG
 - Reproduces the μ-mag output orders of magnitude faster
 - · Generates waveforms fast enough for channel simulation.
- Convolve the output with the RHS, slice → readback
- SNR, BER, SFR (sector failure rate) can be estimated
 - Varying media parameters (K_u, M_s)
 - Varying head parameters (write head field, RHS)
 - Varying writing parameters (eg: TP, BL)



GFP - Model order reduction (MOR)

- Numerator and denominator are multi-dimensional arrays.
- Arrays can be summed over one of their dimensions
 - · Reduces the array dimensionality
 - Increases samples per bin \rightarrow reduces noise
 - · Takes out effect of the removed variable



Elidrissi et. al "Modeling of Two-Dimensional Magnetic Recording and a Comparison of Data Detection Schemes", IEEE Trans. Magn., pp 3685 - 3690, Vol 47, No. 10, Oct 2011

GFP - Density estimation

- GFP channel simulations were run with state-of-the-art HDD parameters (for 2011)
- CBL of 10nm, 12nm, 13nm and 15nm are used here



Elidrissi et. al "Modeling of Two-Dimensional Magnetic Recording and a Comparison of Data Detection Schemes", IEEE Trans. Magn., pp 3685 - 3690, Vol 47, No. 10, Oct 2011

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The Equalizer



- The equalizer shapes the channel response into something more suitable for the detector
- · Unshaped channel response is typically fairly long.
- \rightarrow Excessive states in the ensuing trellis-based detector.
- We design the equalizer such that the combined channel is of some known, more manageable shape.
- Types of equalization:
 - Full response equalization
 - Partial response equalization
 - Generalized partial response (GPR)



Equalizer: Full response equalizer

Full response equalizer attempts to completely invert the channel response:



Equalizer designed to minimize the power of e_k : $E[e_k^2]$ This is the MMSE criterion

• If the equalizer is properly designed, detection of the input bits becomes easy: just threshold.



Problem with full response equalizer: Noise enhancement
 FR Equalizer: W(f)



Generalized/Partial Response Equalizer Design

- Partial Response Equalizer Design: solve w given g and h.
- · Generalized partial response (GPR): solve w and g given h.



$$y_k = \mathbf{w}^T \mathbf{r}_k$$
, where $\mathbf{w} = [w_0 w_1 \dots w_{N_w-1}]^T$ and $\mathbf{r}_k = [r_k r_{k-1} \dots r_{k-N_w+1}]^T$
 $d_k = \mathbf{g}^T \mathbf{a}_k$, where $\mathbf{g} = [g_0 g_1 \dots g_{N_g-1}]^T$ and $\mathbf{a}_k = [a_k a_{k-1} \dots a_{k-N_g+1}]^T$
 $\mathbf{E}[e_k^2] = \mathbf{E}[d_k^2 - 2d_k y_k + y_k^2] = \mathbf{g}^T \mathbf{A} \mathbf{g} - 2\mathbf{g}^T \mathbf{P} \mathbf{w} + \mathbf{w}^T \mathbf{R} \mathbf{w}$

where
$$\mathbf{A} = \mathsf{E} \begin{bmatrix} a_{k}a_{k} & a_{k}a_{k-1} & \cdots & a_{k}a_{k-Ng+1} \\ a_{k-1}a_{k} & a_{k-1}a_{k-1} & \cdots & a_{k-1}a_{k-Ng+1} \\ \vdots & \vdots & \ddots & \vdots \\ a_{k-Ng+1}a_{k} & a_{k-Ng+1}a_{k-1} & \cdots & a_{k-Ng+1}a_{k-Ng+1} \end{bmatrix} \Rightarrow N_{g} \times N_{g}$$

 $\mathbf{P} = \mathsf{E} \begin{bmatrix} a_{k}r_{k} & a_{k}r_{k-1} & \cdots & a_{k-Ng+1}a_{k-Ng+1} \\ a_{k-1}r_{k} & a_{k-1}r_{k-1} & \cdots & a_{k-1}r_{k-Nw+1} \\ \vdots & \ddots & \vdots \\ a_{k-Ng+1}r_{k} & a_{k-Ng+1}r_{k-1} & \cdots & a_{k-Ng+1}r_{k-Nw+1} \\ \vdots & \ddots & \vdots \\ r_{k-1}r_{k}r_{k} & r_{k}r_{k-1} & \cdots & r_{k}r_{k-Nw+1} \\ r_{k-1}r_{k} & r_{k-1}r_{k-1} & \cdots & r_{k-1}r_{k-Nw+1} \\ \vdots & \ddots & \vdots \\ r_{k-Nw+1}r_{k} & r_{k-Nw+1}r_{k-1} & \cdots & r_{k-Nw+1}r_{k-Nw+1} \end{bmatrix} \Rightarrow N_{w} \times N_{w}$



Generalized Partial Response Equalizer Design

- Use Lagrange multipliers to perform constrained minimization
 - Force $g_0 = 1 \Rightarrow [100...0]\mathbf{g} = \mathbf{i}^T \mathbf{g} = 1$
 - To avoid the trivial solution g=w=0
- The method of Lagrange multipliers defines a cost function J
 - Things we want to minimize plus the Lagrange multiplier $\boldsymbol{\lambda}$
 - Differentiate J wrt to the minimization variables and λ
 - Solve the ensuing equations

Cost function:
$$J = \mathbf{g}^T \mathbf{A} \mathbf{g} - 2\mathbf{g}^T \mathbf{P} \mathbf{w} + \mathbf{w}^T \mathbf{R} \mathbf{w} - 2\lambda (\mathbf{i}^T \mathbf{g} - 1)$$

 $\nabla_g J = 2\mathbf{A} \mathbf{g} - 2\mathbf{P} \mathbf{w} - 2\lambda \mathbf{i} = \mathbf{0}$ (1)
 $\nabla_w J = -2\mathbf{P}^T \mathbf{g} + 2\mathbf{R} \mathbf{w} = \mathbf{0}$ (2)
 $\nabla_\lambda J = \mathbf{i}^T \mathbf{g} - 1 = \mathbf{0}$ (3)

$$(2) \Rightarrow \mathbf{w} = \mathbf{R}^{-1} \mathbf{P}^{T} \mathbf{g}$$

$$(1) \Rightarrow 2\mathbf{A}\mathbf{g} - 2\mathbf{P}\mathbf{R}^{-1} \mathbf{P}^{T} \mathbf{g} - 2\lambda \mathbf{i} = 0 \Rightarrow \mathbf{g} = \lambda \left(\mathbf{A} - \mathbf{P}\mathbf{R}^{-1} \mathbf{P}^{T}\right)^{-1} \mathbf{i}$$

$$(3) \Rightarrow \lambda \mathbf{i}^{T} \left(\mathbf{A} - \mathbf{P}\mathbf{R}^{-1} \mathbf{P}^{T}\right)^{-1} \mathbf{i} = 1 \Rightarrow \lambda = \frac{1}{\mathbf{i}^{T} \left(\mathbf{A} - \mathbf{P}\mathbf{R}^{-1} \mathbf{P}^{T}\right)^{-1} \mathbf{i}}$$



The iterative detector



- After equalization to known target, iterative detector returns detected/decoded bits
- · The iterative detector consists of:
 - Data-dependent noise-predictive (DDNP) detector
 - The noise from the equalized channel is coloured and data-dependent
 - · White-noise Viterbi Algorithm is suboptimal
 - DDNP Viterbi Algorithm performs better.
 - Optimal if the noise is as expected by the model (AR)
 - LDPC decoder
 - The sum-product algorithm (SPA)
 - Operating on a factor graph (FG)



DDNP Model



"A Signal Dependent Auto-regressive Channel Model", A. Kavcic, A. Patapoutian, IEEE Trans. Magn. Sept 1999

- The noise in the model is auto-regressive (AR) of order L
 - · The AR coefficients are data-dependent
 - · The noise coloration is data-dependent
- The signal component comes from a look-up table
 - Signal vector α_k defined by range $I_1 \rightarrow I_2$



DDNP Noise prediction



- Noise coloring filter
- Auto-regressive
- · White noise input

$$n_k = \sum_{l=1}^L b_l n_{k-l} + \nu_k$$

- Noise whitening filter
- Moving average
- Inverse of noise coloring filter

$$\nu_k = n_k + \sum_{l=1}^L -b_l n_{k-l}$$

- Whitened noise ν_k less power than colored noise n_k
- DDNP detector uses ν_k to compute metrics in the trellis.





DDNP performance comparison via GFP model



- · Results over the GFP channel model at various bit lengths
- Using iterative detector, with and without DDNP.
- DDNP always performs better than non-DDNP
- Shorter BL → more media noise → more gain for DDNP



Iterative detector



- Detector algorithm (DDNP-SOVA) operates on the trellis
 - Returns channel probabilities, passes the decoder.
- Decoder algorithm (SPA) operates on the factor graph
 - · Iterates a number of times between bit and check nodes
 - Returns a-priori probabilities to the detector
- · Global iterations from detector trellis to decoder FG.
- Sometimes an interleaver sits between the two.



The joint Viterbi detector/decoder



- State of the art: Iterative detector
 - SOVA detector knows about the channel
 - SPA decoder knows about the code
 - Exchange information between them iteratively.
- JVDD: proposed competitor to the iterative detector
 - Knows about both the channel and the code
 - Performs both detection and decoding jointly on a trellis
 - · Based on the Viterbi algorithm
 - Conditionally optimal over AWGN/ISI channel
 - · Optimum with sufficient computing resources
 - Main challenge:
 - · Managing the complexity/performance trade-off







JVDD Codes

Random LDPC code



- Two H matrices shown with Black pixels = 1, white pixels = 0
 - Top figure shows a (512,52) random code.
 - Bottom figure shows a (512,52) JVDD code.
- Last "1" in each row of both matrices marked by red pixel.
 - Random LDPC code has last "1" in each row clumped towards end of matrix
 - JVDD code has last "1" in each row more uniformly distributed through the matrix
- Parity checking in JVDD can occur throughout trellis when using JVDD codes



JVDD over AWGN/ISI Channel: short CWL



- Performance plots for the JVDD vs iterative detector
- · Complexity plots (in avg no of surv) for the JVDD.
- Both iterative detector and JVDD using random codes
- JVDD greatly outperforming iterative detector at short CWL
 - By more than 2dB.



JVDD over AWGN/ISI Channel: inter CWL



- Iterative detector using both random codes and JVDD codes
- JVDD using only JVDD codes.
- Iterative doing better with random than JVDD codes.
- But even with random codes, JVDD outperforms it with JVDD codes.



JVDD over AWGN/ISI Channel: long CWL



- JVDD codes can't handle⇒ introduce maxNoSurv parameter.
 - · Limit maximum number of surv in trellis
- JVDD codes outperform random codes (for JVDD algo) even with maxNoSurv parameter.
- Iterative performance begins to surpass JVDD at long CWI



JVDD over a magnetic recording channel

Constant Media Parameters

91	Ms	480 emu/cc	
AC	σ_{Ms}	3%	
4	Ku	3.6e6 erg/cc	
1024	σ_{Ku}	3%	
15	A _x	3e-7 erg/cm	
6	σ_{Ax}	3%	
7	easy axis	0°	
20%	easy axis σ	1.7°	
30			
	AC 4 1024 15 6 7 20%	AC σ_{Ms} 4 K_u 1024 σ_{Ku} 15 A_x 6 σ_{Ax} 7easy axis20%easy axis σ	

Variable Media Parameters

Track pitch TP (nm)	30	34	42	50
Reader pitch	10	26	34	42
RP(nm)	48	50		
Reader offset RO (nm)	-4	-2	-0	
Equalizer type	1D	SMR	TDMR	
Reader Config- uration	single	triple		

Detector/Decoder Parameters

JVD	D threshold	2	Iterative	LDPC iterations	150
	maxSurv	10000, 30000		global iterations	6
	code	GDLD		code	random LDPC
	CWL	512,1024,2048		CWL	512,1024,2048

Sample μ -mag simulation output: DC-erase, TP=50 4 tracks x 40bits





JVDD over a magnetic recording channel



- JVDD generally outperforming iterative detector
- Improvement gap decreasing with increasing CWL
- · But benefits at increasing CWL include:
 - Waterfall curves become more steep
 - Code rate increases
- Reducing TP ⇒ more noise from track-edge ⇒ more errors ⇒ lower code rates needed to correct them.



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Future HDD Technologies



Current Technology: Perpendicular Recording

Upcoming Technology: 2D/shingled magnetic recording



Future Technology: MAMR (µ-wave Assisted Magnetic Recording

Future Technology: HAMR (Heat Assisted Magnetic Recording

Future Technology: BPMR (Bit Patterned Media Recording





Longitudinal recording writes bits in the plane of the medium using fringe field from head.

Perpendicular recording changed the orientation of the bits perpendicular to the medium and included an SUL. Medium now in the head gap.

SMR/TDMR overlaps adjacent tracks into a 2D block of bits. 2D Processing also occurs using 2D equalization, detection and decoding techniques.

MAMR uses microwave frequencies to excite the grain magnetizations making them easier to switch. Enable switching smaller higher Ku grains.

HAMR uses heat to thermally excite the grains and reduce Ku. This again enable smaller thermally stable grains to be used.

BPMR writes 1 channel bit per magnetic particle (island). These islands are engineered on a lattice that enables them to store data more reliably.



SMR/TDMR



- SMR: Shingled Magnetic Recording
 - Enables writing of narrower tracks with wider writers
 - Writers designed for tight gradients on trailing edge, and one side
 - Information written in blocks rather than tracks
 - SMR considered an intermediate technology to carry the industry over until MAMR/HAMR/BPMR mature
 - Is currently being implemented in product today.



SMR/TDMR



- TDMR: Two-dimensional Magnetic Recording
 - Enables reading of narrower tracks with wider readers
 - Wider readers⇒ ITI (intertrack interference)
 - TDMR mitigates ITI by using 2D detectors
 - · Requires multiple readback to be processed at a go
 - · Involves putting multiple reader elements on the head
 - Involves changing 1D channel \Rightarrow 2D channel
 - 2D channel
 - Greater computational complexity
 - Best TDMR scheme still being investigated



Energy assisted Magnetic Recording: MAMR



- Energy assisted magnetic recording helps to overcome the superparamagnetic limit
- Use small grains with higher $K_u \Rightarrow$ difficult to write
- Inject energy (heat or microwaves) during writing.
 - · Energy helps the switching process
 - Effectively brings down K_u during writing
- · Limiting factor:
 - How well the energy can reduce K_u
 - The combined gradient of the magnetic + energy assist



Energy assisted Magnetic Recording: MAMR



- The energy source in MAMR is microwaves
 - · Grains have a natural tendency to precess
 - When excited near resonant frequency ⇒ become unstable
 - · Application of a magnetic field completes the switch
- Microwave field applied via spin torque oscillator (STO)
 - STO oscillates when driven with a current
- Challenges:
 - Generating a stable μ -wave assisting field
 - Design/placement of the STO.



Energy assisted Magnetic Recording : HAMR



- · Energy source in HAMR is heat from a laser.
 - Grains lose their magnetism when hot $(T > T_c, T_c \gtrsim 600 \text{K})$
 - · Cooling grains align with external field.
- Challenges
 - · Generting a small thermal spot (The NFT)
 - Light delivery to the NFT
 - Creating very small high K_u grains
 - Overcoat and lubricants that work at high temp



BPMR



- BPMR aligns the "grains" on the medium in a regular lattice.
- Ordered rather than random grains. SNR ≁ log(N).
- Store each bit on 1-2 islands (7-12 grains for granular media)
- Challenges
 - Creation of regular arrays of nanoparticles for the media
 - Electron lithography
 - Cheaply manufacturing BPMR disks
- Synchronizing the write field to the islands



